

# Intermittent structures in solar wind turbulence from MHD to sub-ion scales at 0.17 AU from the Sun

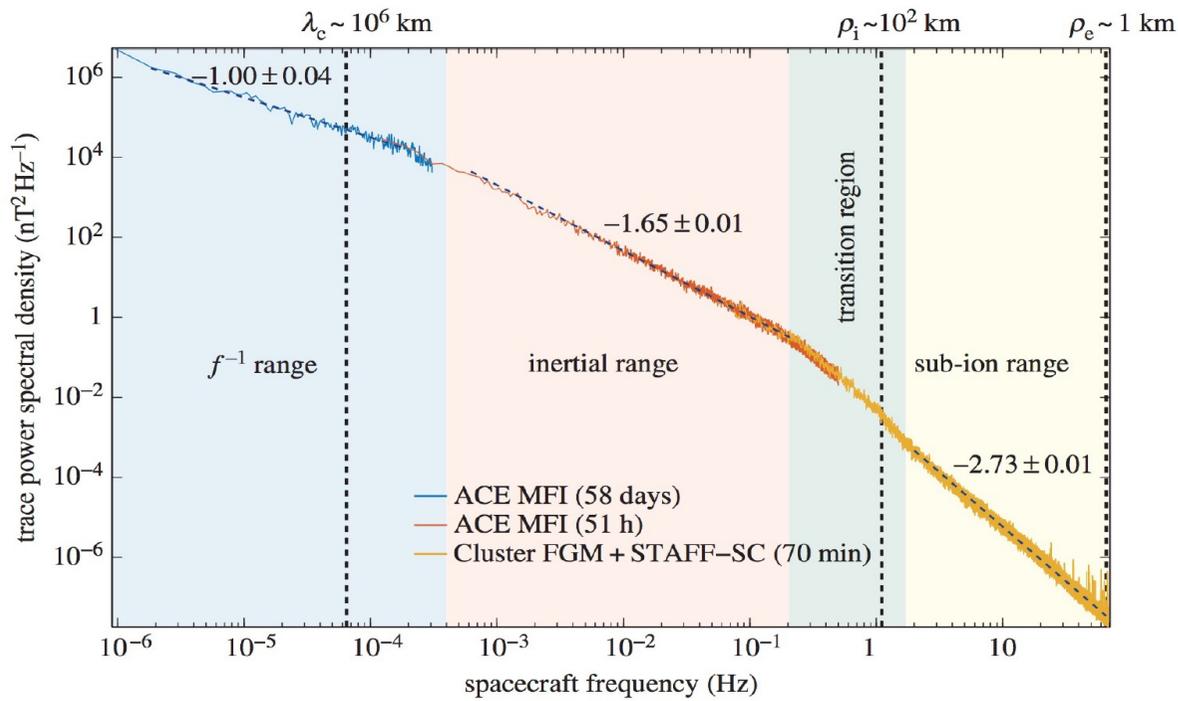
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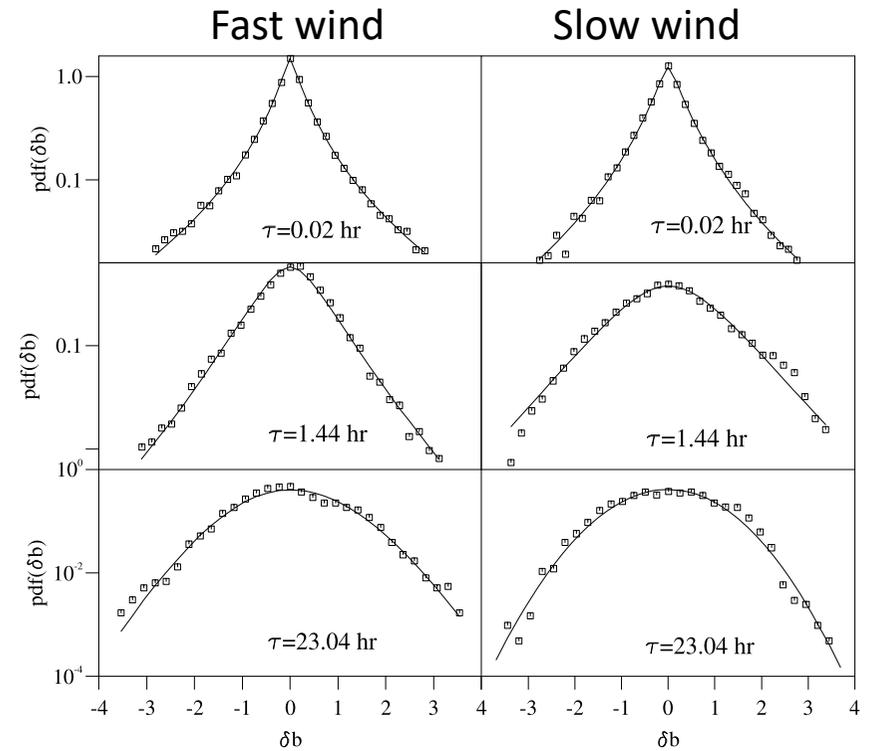
23-05-2022

# Solar wind turbulence

[Kiyani et al. 2015]

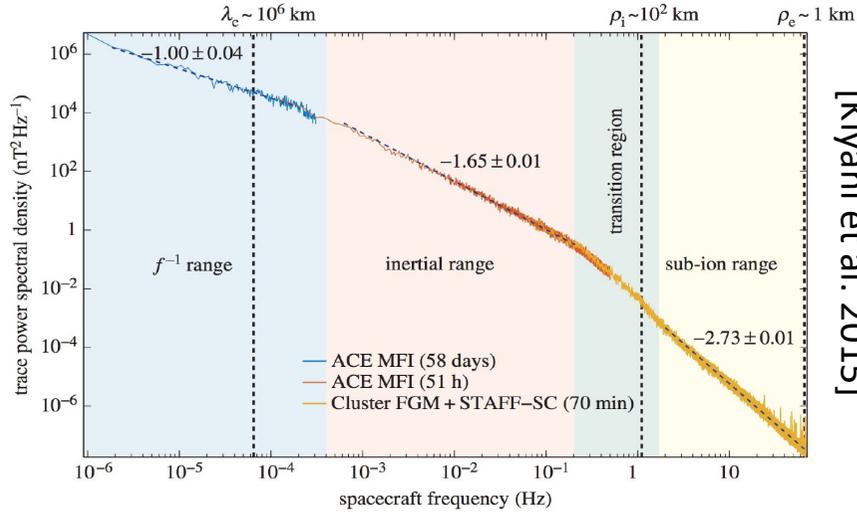


[Sorriso-Valvo et al. 1999]



$$|\Delta \mathbf{b}| = |\mathbf{b}(s + \Delta s) - \mathbf{b}(s)|$$

# Structures in the solar wind

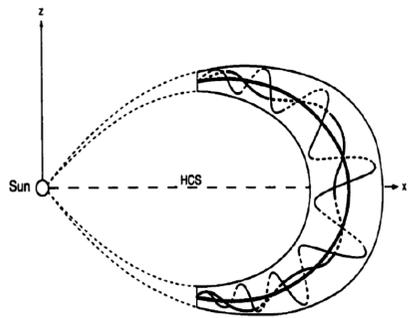


[Kiyani et al. 2015]

- F-1 range: Flux ropes [Zhao 2020], [Janvier 2014]  
Flux tubes [Borovsky 2008]
- Inertial range: Discontinuities [Knetter 2004],  
Magnetic holes [Stewens 2007], [Karlsson 2021]  
Alfvén vortex [Lion 2016]
- Ion scales: Discontinuities, shocks, magnetic holes,  
Alfvén vortices [Perrone 2016, 2017]
- Sub-ion range: Discontinuities [Perri 2012], [Greco 2016]  
Magnetic hole [Liu 2019]

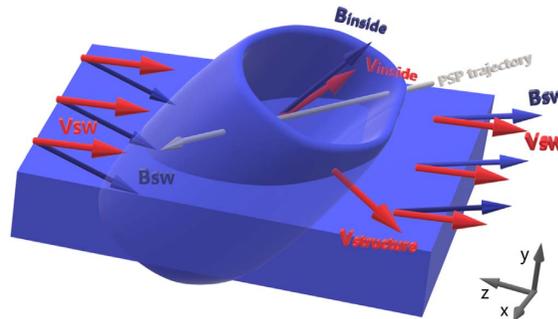
Is there a relation between all these structures across the cascade?

F<sup>-1</sup> range  
Flux rope



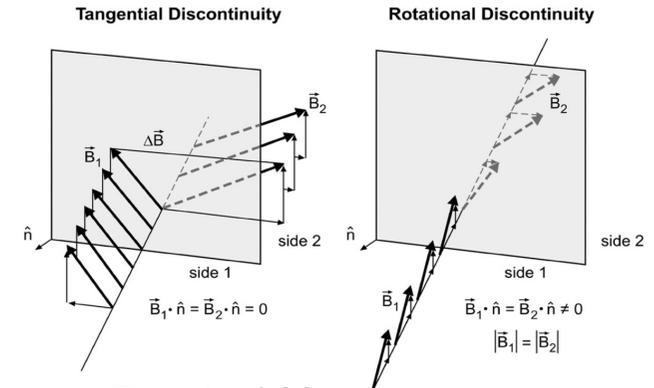
Burlaga 1990

Inertial range  
Alfvénic-type magnetic structure



Krasnoselskikh 2020

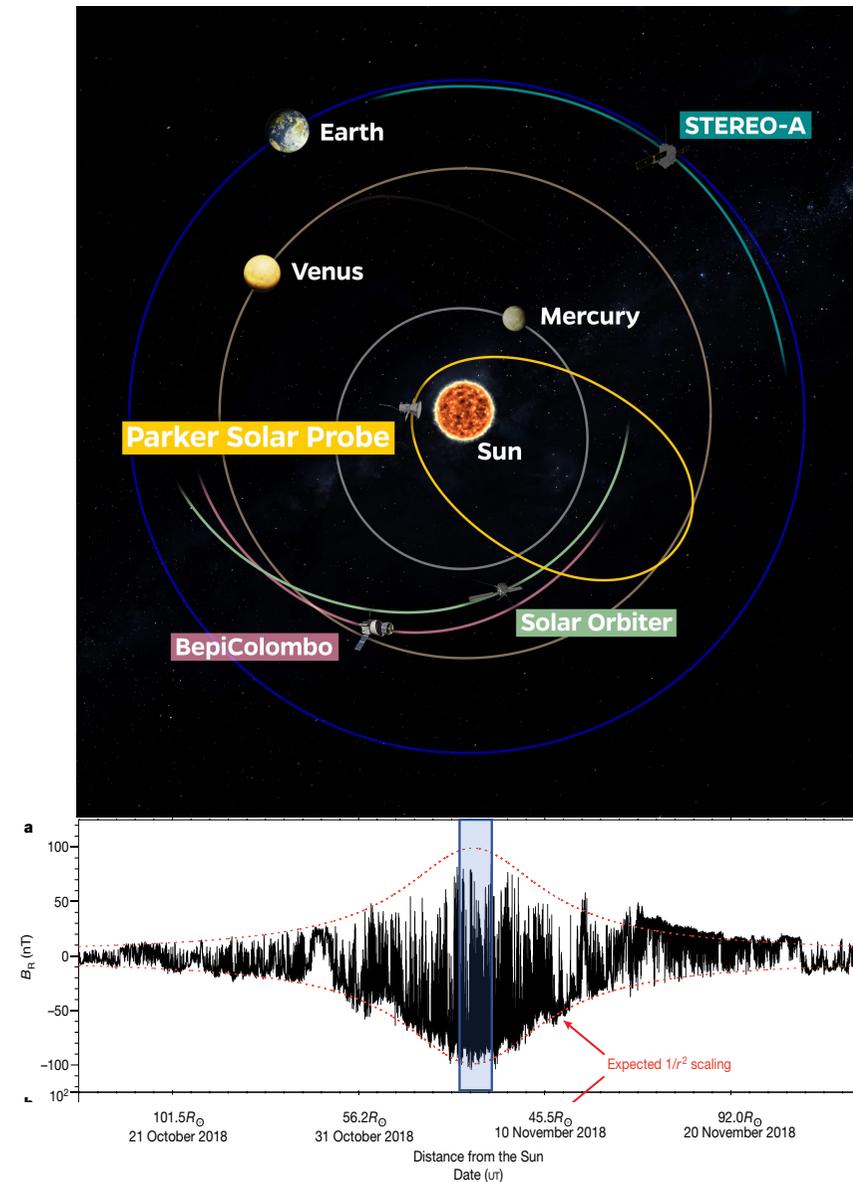
Ion-scale  
discontinuities



Tsurutani 2011

# Parker Solar Probe First encounter: 2018-11-06

- Unique opportunity exploring the turbulence in the “young” solar wind  
Radial distance  $R \sim 25 \cdot 10^9 \text{ m} \sim 0.17 \text{ au}$
- The main goal of our investigation is to characterize solar wind coherent structures from MHD to sub-ion scales



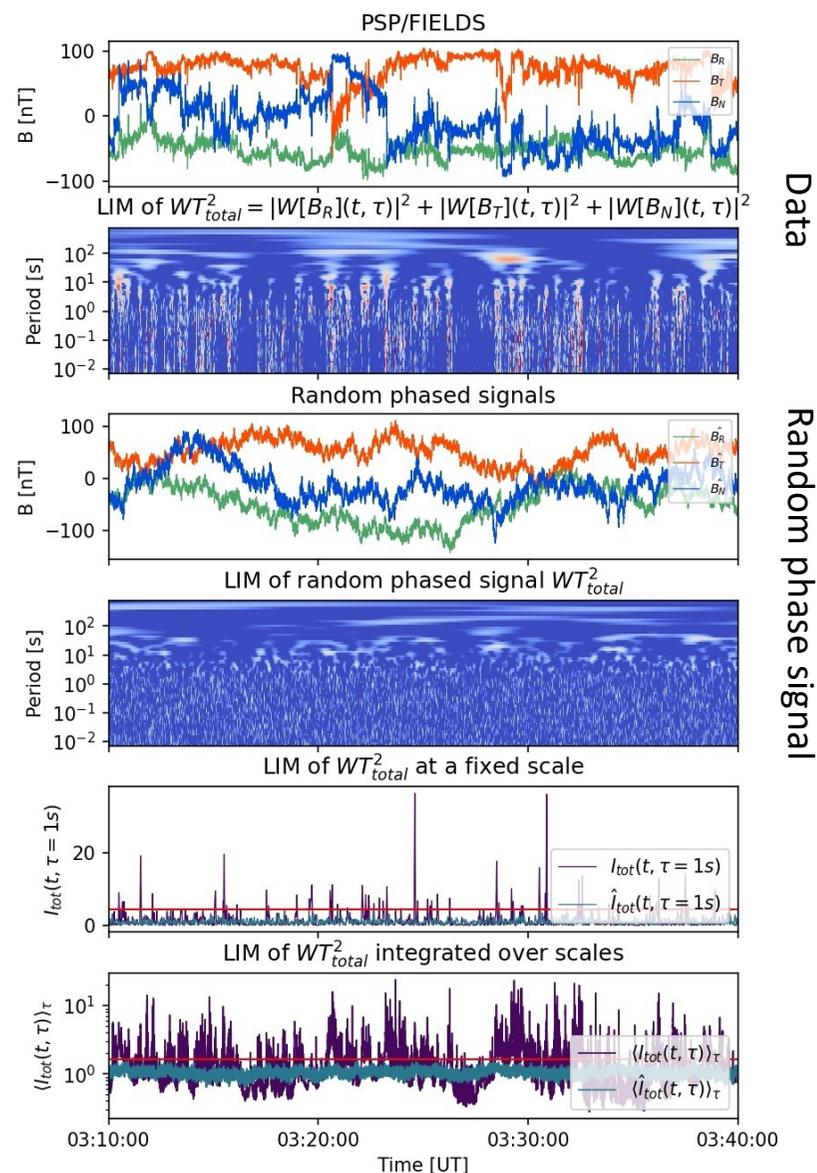
Bale et al., 2019, Nature

# Detection of coherent structures at 0.17 AU

- Morlet Wavelet
- Total local intermittency measure  $I(t, \tau)_{tot}$  shows the relative total energy of fluctuations at a given moment in time at a given scale  $\tau$

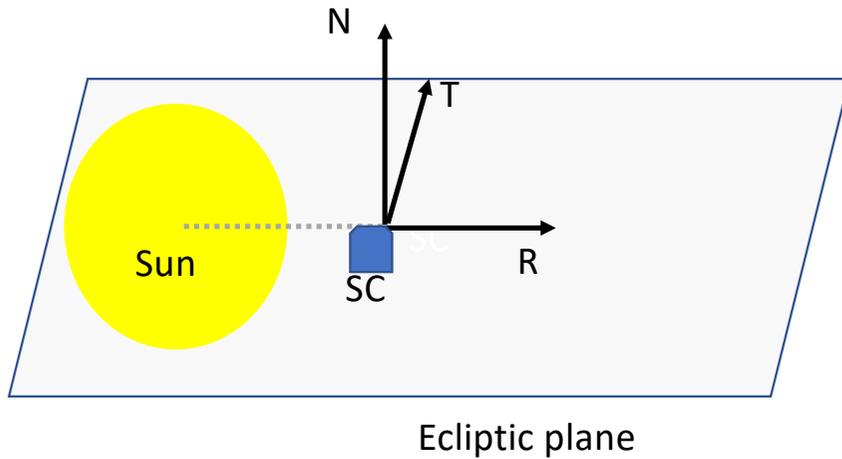
$$I(t, \tau)_{tot} = \frac{\sum_{i=R,T,N} |W[B_i](t, \tau)|^2}{\langle \sum_{i=R,T,N} |W[B_i](t, \tau)|^2 \rangle_{t \in T'}}$$

- To detect the structures we compare magnetic field measurements with an random-phased signal
- **Vertical lines: coupled phases across scales -> Coherent structures**

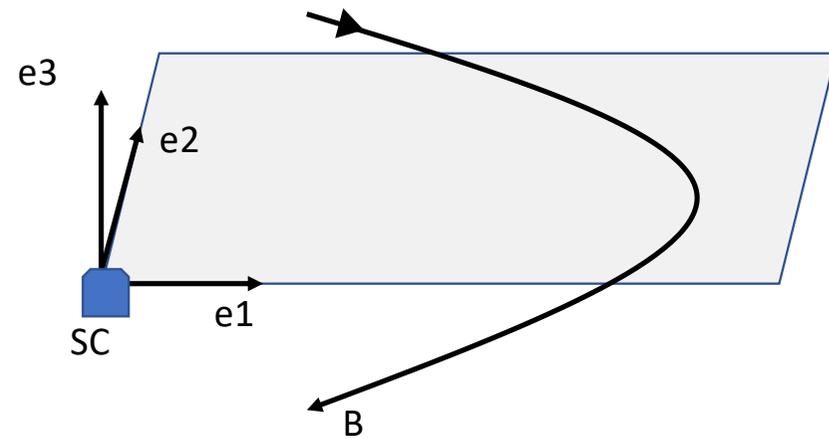


# Reference frames

Radial Tangential Normal  
(RTN) reference frame



Local Minimum variance  
(MVA) Reference frame



$\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3$  are eigenvectors of the magnetic field covariation matrix

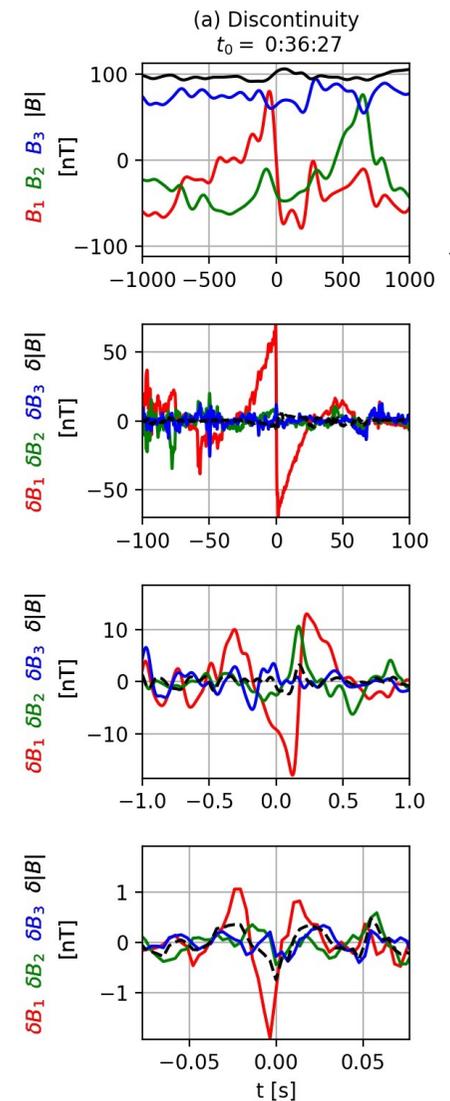
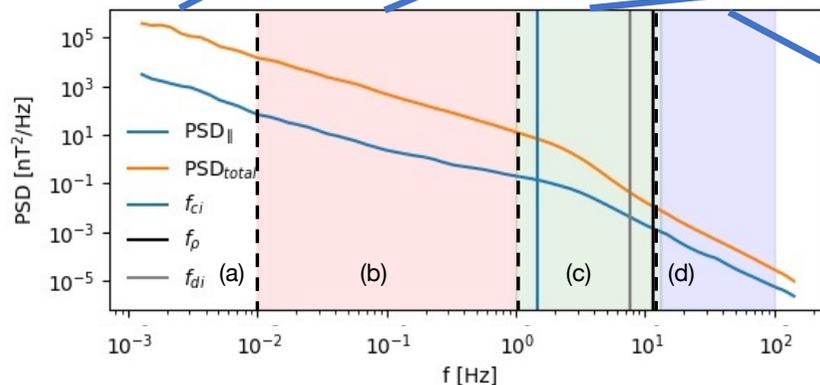
Relations of the corresponding eigenvalues  $\lambda_2/\lambda_1, \lambda_3/\lambda_2$  show if  $\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3$  directions are well-defined

# Examples of coherent structures

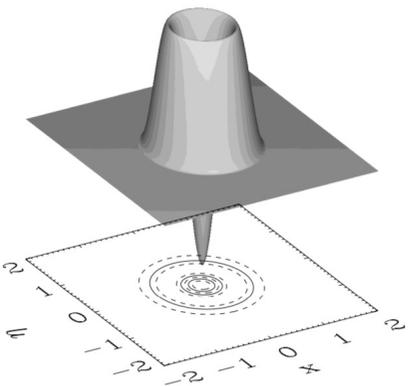
## 1. Planar discontinuity

- Current sheet at MHD scales
  - Sub-ion scales: vortex-like structure?
- Is it formed as a result of current sheet instability/reconnection?

To define fluctuations within a frequency range, we apply a band-pass filter  $f \in F_j$

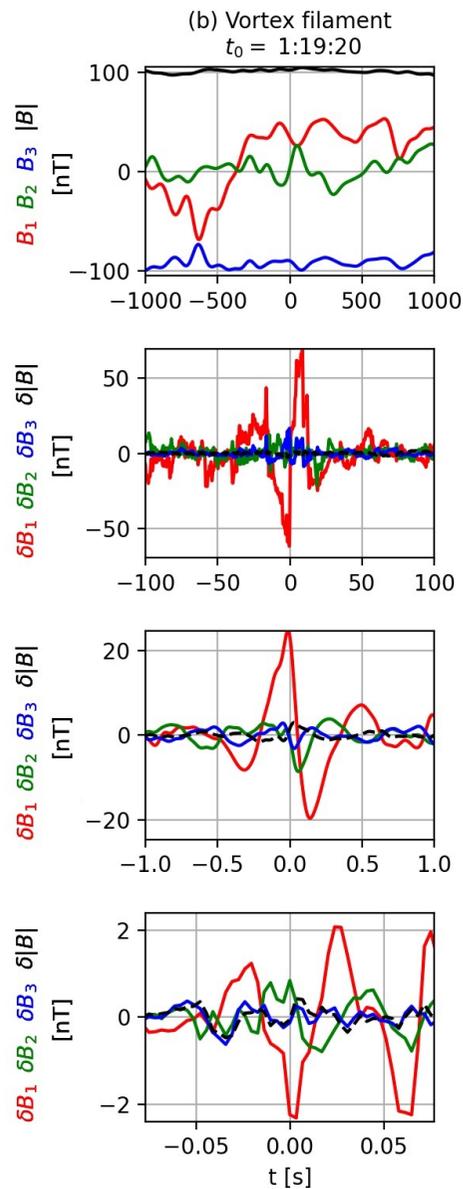
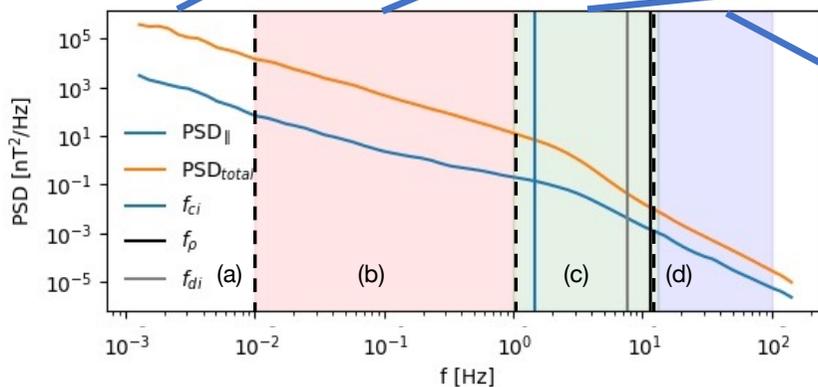


# 2. Alfvén vortex like event



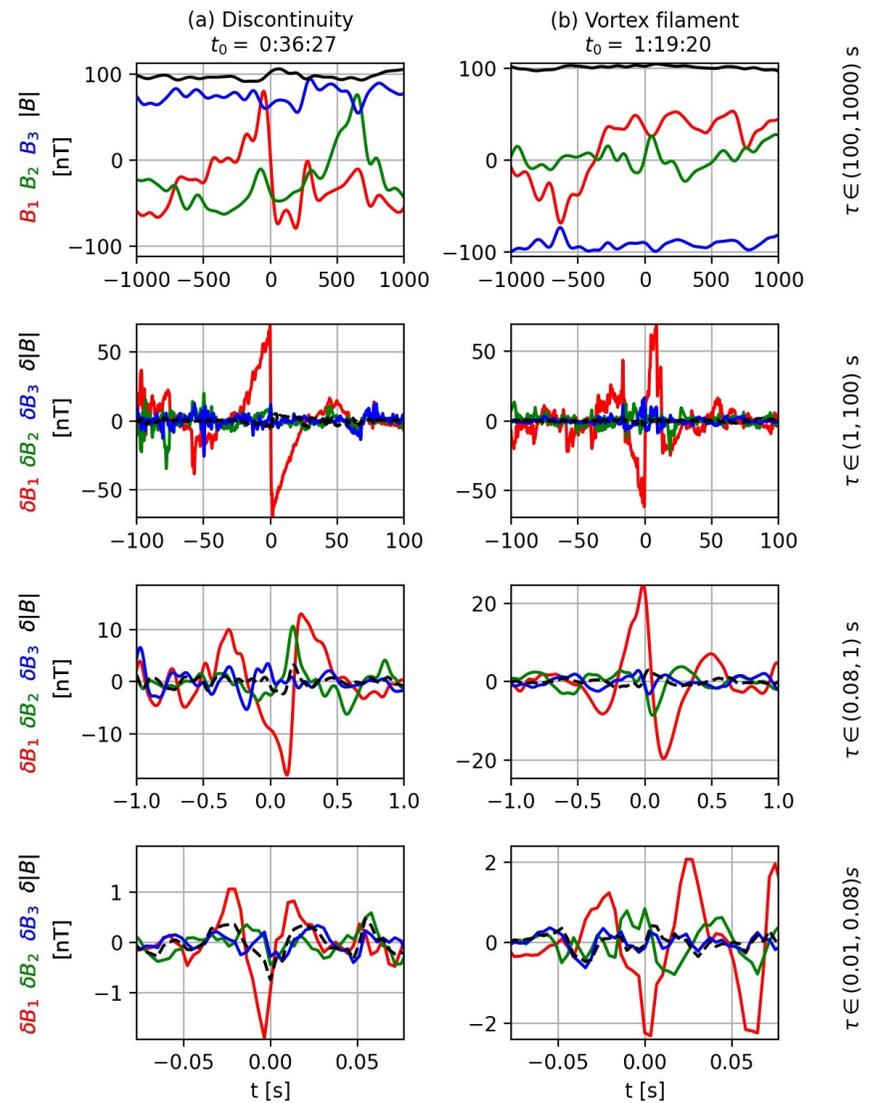
- Incompressible MHD vortex [Petviashvili & Pokhotelov 1992]
- Compressible MHD + Ion-scales vortex model: Jovanovic et al. 2020;

Figure from Alexandrova 2008

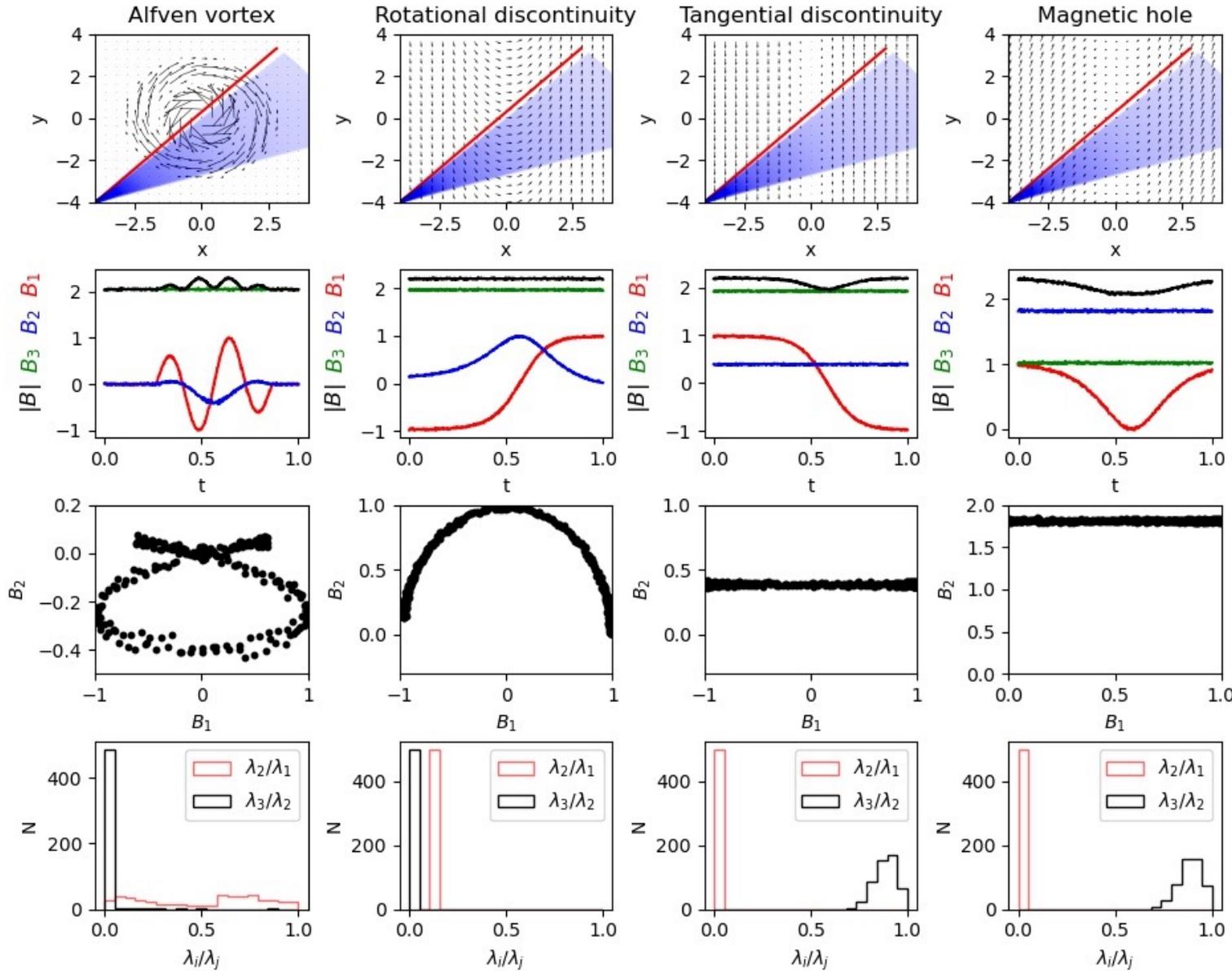


# Two examples of coherent structures across cascade

- Small scale events are substructure of large scales ?
- We observe a large number of small scale events within a large scale structure.
- Topology of the observed events depends strongly on the satellite trajectory.



# Simulation of coherent structures crossings



$\lambda_1$  – maximum variation  
 $\lambda_2$  – intermediate variation  
 $\lambda_3$  – minimum variation

Dipole Vortex  $\sim$  RD

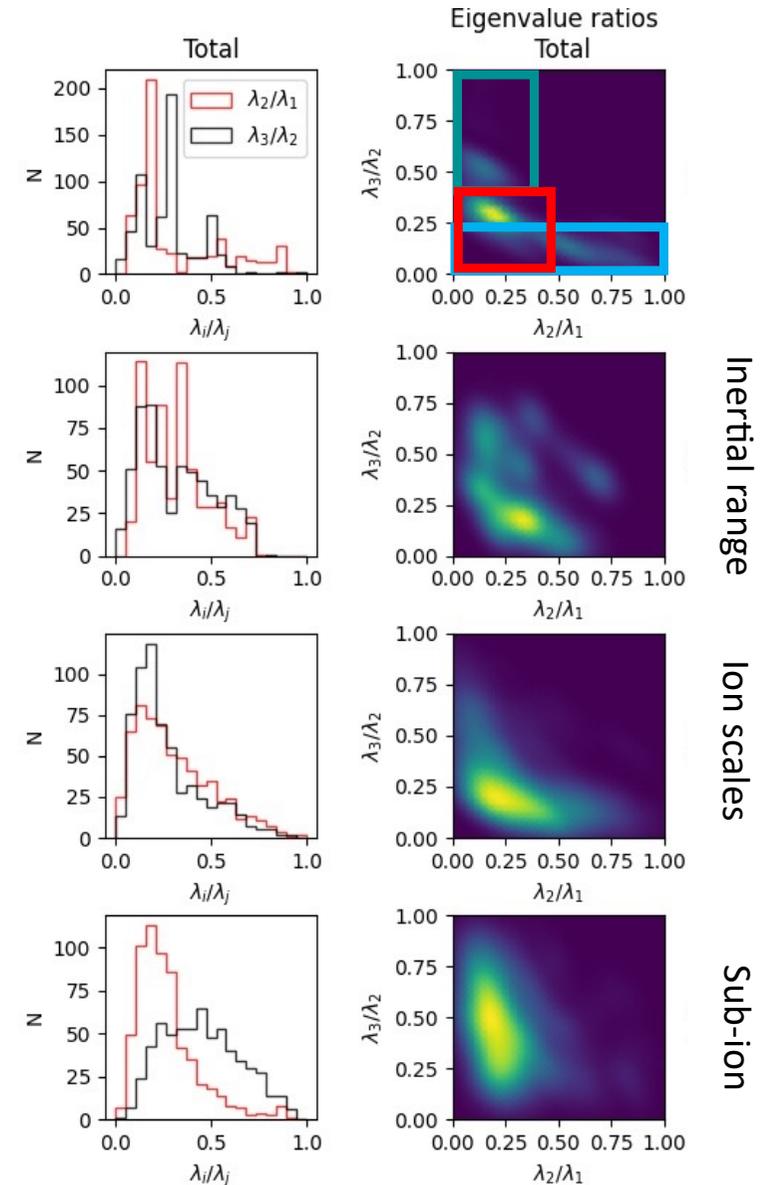
# Statistics: 600 structures for 6h of observation

MVA eigenvalue ratios  $\lambda_2/\lambda_1$ ,  $\lambda_3/\lambda_2$  allow to distinguish between different types of structures:

- **Upper left area** – tangential discontinuities/magnetic holes
- **Lower elongated rectangle** – monopole Alfvén vortices
- **Area at the zero vicinity** – rotational discontinuities/dipole vortex

So we have:

- MHD: rotational discontinuities and vortices dominate
- ion scales: RD & increasing population of vortices
- The population of tangential discontinuities/magnetic holes increases towards the smallest scales.



# Conclusion

- We observe solar wind coherent structures from MHD to ion kinetic scales and below.
- We try to determine topology of the structures using minimal variance analysis applied on the data and on the model structures.
- We observe evolution of the topology of magnetic fluctuations from MHD to sub-ion scales.

Range of scales	Possible dominant type of structures
<b>MHD Inertial range</b>	<b>Rotational discontinuities/Dipole vortex</b>
<b>Ion scales</b>	<b>Alfven vortexes &amp; rotational discontinuities</b>
<b>Sub-ion scales</b>	<b>Tangential discontinuities/ Magnetic holes</b>